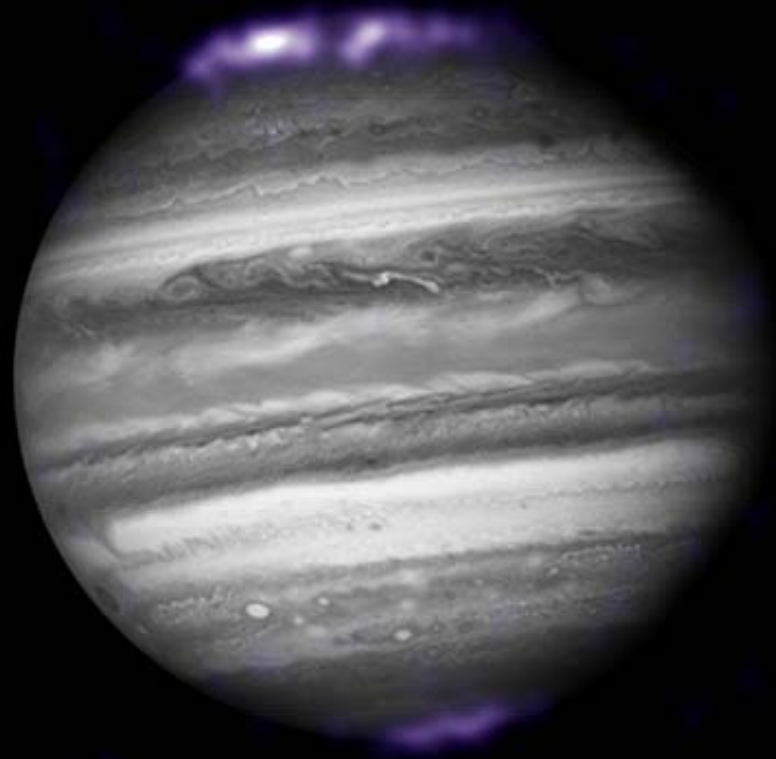


Magnetic Reconnection and the Aurora

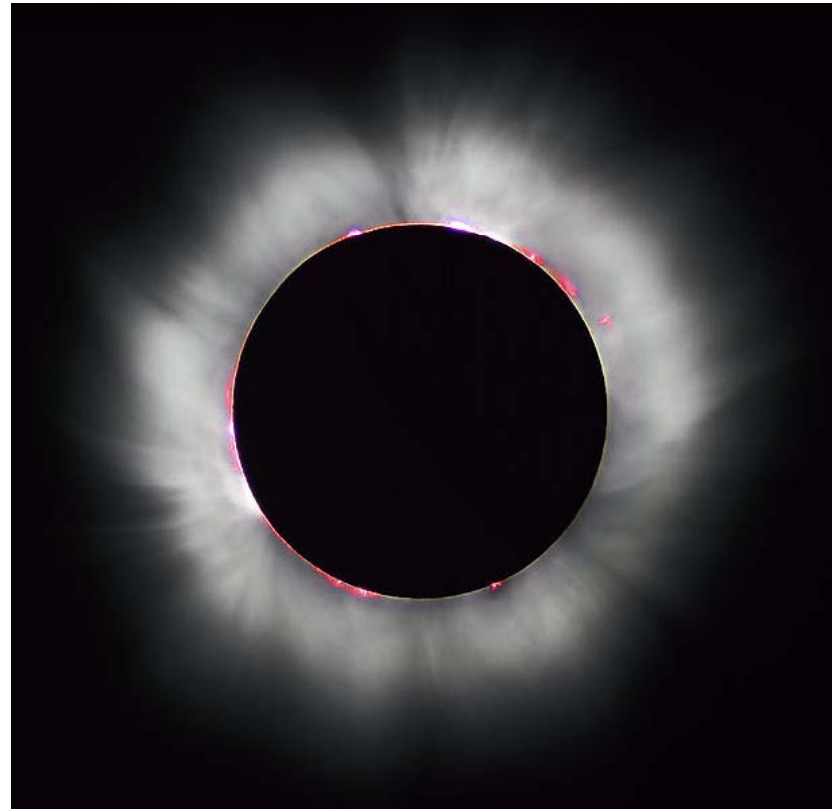


Outline

- Introduction
- Frozen-in Field Lines
- Field Line Topology
- Non-Ideal Effects/Magnetic Reconnection
- Conclusion

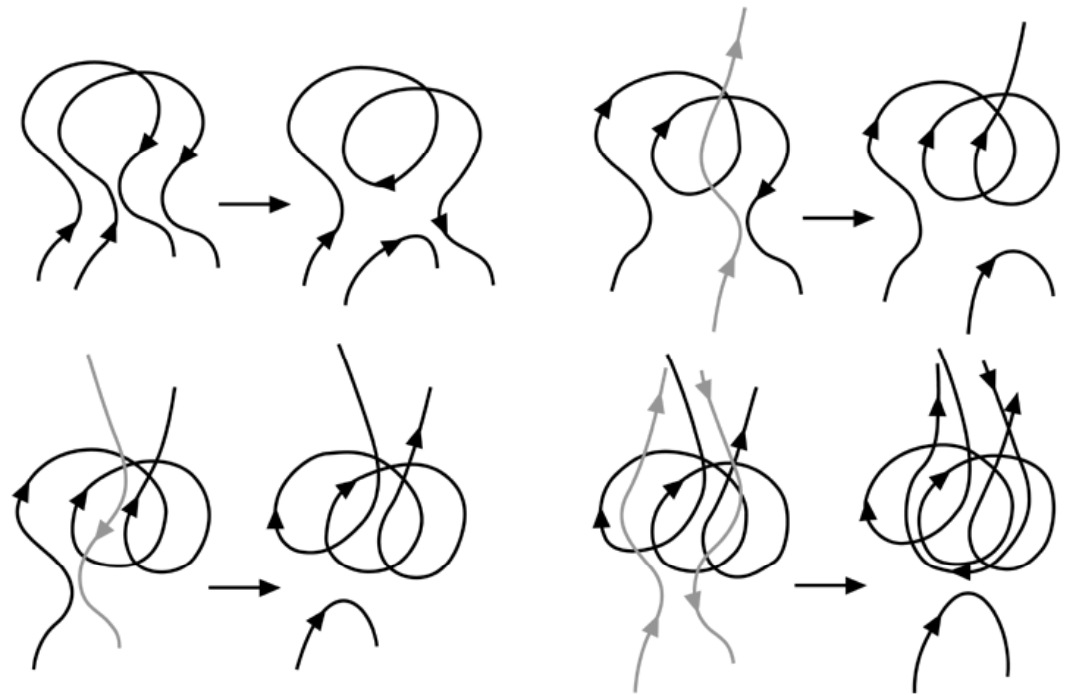
Interplanetary Plasma

- Corona of the sun
- Mostly Ionized Hydrogen
- Carries “Frozen-in” Magnetic Field

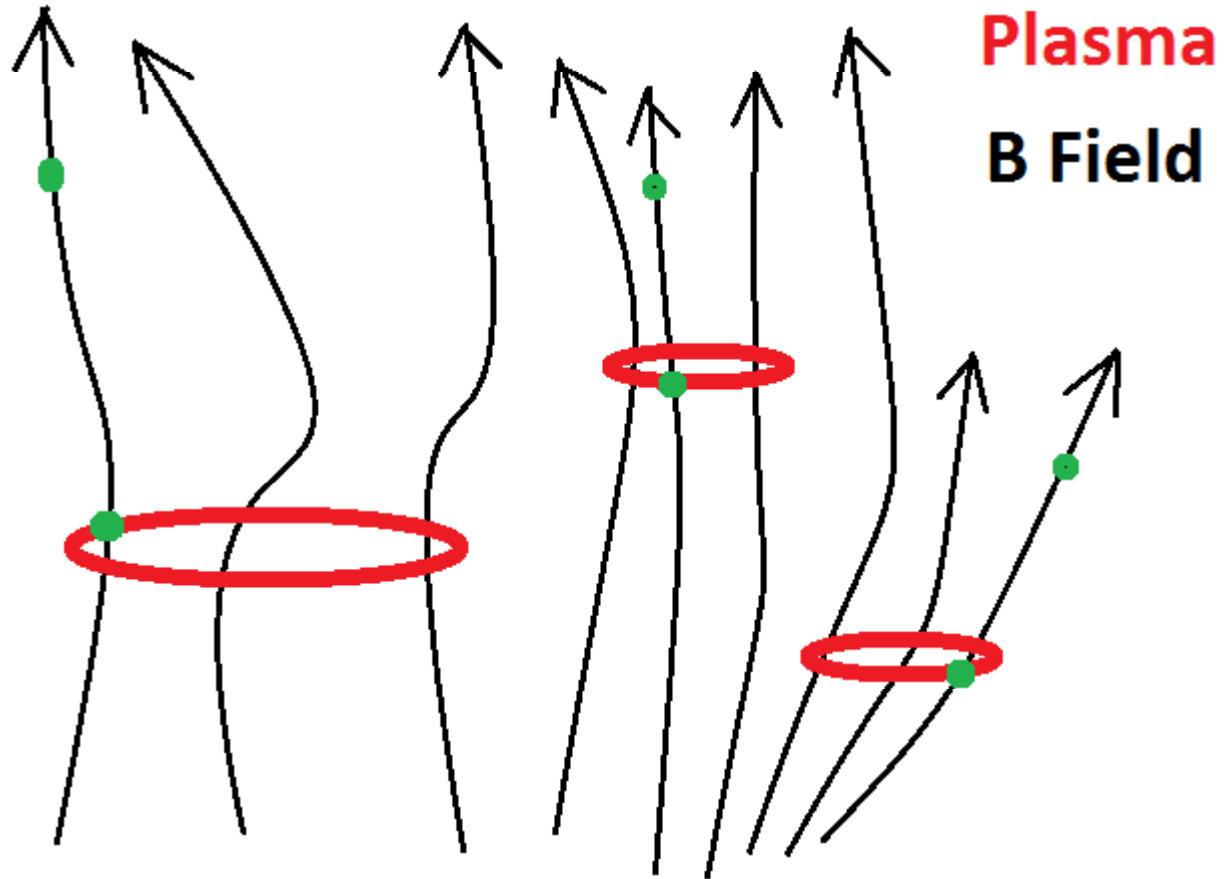


Magnetic Reconnection

- Auroral Trigger Mechanism
- Release of stored Magnetic Energy
- Conversion to Kinetic Energy
- Return to a Relaxed State



Frozen-In Fields



Magnetic Flux Tube Conservation

Generalized Ohm's Law (c=1) $E + \mathbf{v} \times \mathbf{B} = \mathbf{R}$

Ideal Approximation $E + \mathbf{v} \times \mathbf{B} = 0$

Faraday's Law $\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B})$

Magnetic Flux $\Phi = \int_S \mathbf{B} \cdot d\mathbf{S}$

Change in Flux $\frac{d\Phi}{dt} = \int_S \frac{\partial \mathbf{B}}{\partial t} \cdot d\mathbf{S} + \oint_{\partial S} \mathbf{B} \cdot (\mathbf{v} \times d\mathbf{l})$

but $\int_S \frac{\partial \mathbf{B}}{\partial t} \cdot d\mathbf{S} = \oint_{\partial S} (\mathbf{v} \times \mathbf{B}) \cdot d\mathbf{l} = - \oint_{\partial S} \mathbf{B} \cdot (\mathbf{v} \times d\mathbf{l})$

$$\Rightarrow \frac{d\Phi}{dt} = 0$$

Field Lines are "Frozen-In" to the plasma.

Conservation of Magnetic Topology

Consider two generating functions for the field lines

$$F_B(\mathbf{x}, s, t) \quad \text{and} \quad F_u(\mathbf{x}, t, t_0)$$

with the following properties

$$\partial_s F_B(\mathbf{x}, s, t) = \mathbf{B}(\mathbf{x}, t) \quad \text{with} \quad F_B(\mathbf{x}, 0, t) = \mathbf{x}$$

$$\partial_t F_u(\mathbf{x}, t, t_0) = \mathbf{u}(\mathbf{x}, t) \quad \text{with} \quad F_u(\mathbf{x}, t_0, t_0) = \mathbf{x}$$

Let a field line connect two points \mathbf{x} and \mathbf{y} at time t_0

$$\mathbf{y} = F_B(\mathbf{x}, s, t_0) \quad \text{for some } s$$

and let \mathbf{x}' and \mathbf{y}' be the points to which \mathbf{x} and \mathbf{y} are transported after a time $t-t_0$.

$$\mathbf{x}' = F_u(\mathbf{x}, t, t_0)$$

$$\mathbf{y}' = F_u(\mathbf{y}, t, t_0) = F_u(F_B(\mathbf{x}, s, t_0), t, t_0)$$

Magnetic Topology cont'd

The topology of the field lines will be considered conserved if \mathbf{x}' and \mathbf{y}' still lie along the same field line, i.e.

$$\begin{aligned} \mathbf{y}' &= F_B(\mathbf{x}', s', t) = F_B(F_u(\mathbf{x}, t, t_0), s', t) \\ \Rightarrow F_u(F_B(\mathbf{x}', s, t), t, t_0) &= F_B(F_u(\mathbf{x}, t, t_0), s', t) \end{aligned}$$

Taking the second mixed derivative at $s=0$ and $t=t_0$ gives

$$\begin{aligned} \partial_s \partial_t F_u(F_B(\mathbf{x}, s, t_0), t, t_0) &= \partial_t \partial_s F_B(F_u(\mathbf{x}, t, t_0), s', t) \\ &= \partial_s [u(F_B(\mathbf{x}, s, t_0), t_0)]_{s=0} = \partial_t [B(F_u(\mathbf{x}, t, t_0), t) \cdot \partial_s s']_{t=t_0} \end{aligned}$$

$$= \nabla u \cdot \mathbf{B} = \partial_t \mathbf{B} + \mathbf{B} \cdot \nabla u + \mathbf{B} \partial_t \partial_s s'$$

Consequence

The total time derivative of B
is proportional to B .

$$\frac{dB}{dt} = B\sigma + B \cdot \nabla u$$

Magnetic nulls cannot be created, only moved.

Non ideal effects give rise to new
magnetic nulls, i.e. magnetic reconnection.